# 太陽大気・太陽風の形成機構 4-3班(原・鈴木 担当)

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#### 2014年10月20日 (Oct.20th, 2014)

# Energy/Momentum/Mass transfer in the atmosphere

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 Extract the kinetic energy of the surface convective turbulence

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- The energy dissipates at appropriate locations



In Situ Heating in the Corona & Wind

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In Situ Heating in the Corona & Wind

$$= \rho v_{\parallel} \left( v^2/2 + h - GM/r \right) + v_{\parallel} B_{\perp}^2 / 8\pi - B_{\parallel} v_{\perp} B_{\perp} / 4\pi$$

#### Energy/Momentum/Mass transfer in the atmosphere T,

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In Situ Heating in the Corona & Wind

Energy flux =  $\rho v_{\parallel} \left( v^2/2 + h - GM/r \right) + v_{\parallel} B_{\perp}^2 / 8\pi - B_{\parallel} v_{\perp} B_{\perp} / 4\pi$ 

Poynting flux at different locations.  $\Rightarrow$  Heating

## **Energy Transfer**

$$F = \rho v_{\parallel} \left( v^2/2 + h - GM/r \right)$$
$$+ v_{\parallel} B_{\perp}^2 / 8\pi - B_{\parallel} v_{\perp} B_{\perp} / 4\pi$$

at different locations ⇒ Heating

# **Energy Transfer**

Example (Hiller+ 2013)

-1.0 -0.5 0.0 0.5 1.0 1.5

log Frequency v<sub>o</sub> (mHz)

2.0

$$F = \rho v_{\parallel} \left( v^2/2 + h - GM/r \right)$$
  
+ $v_{\parallel} B_{\perp}^2/8\pi - B_{\parallel} v_{\perp} B_{\perp}/4\pi$   
at different locations  
 $\Rightarrow$  Heating

# **Energy Transfer**

$$F = \rho v_{\parallel} \left( v^2/2 + h - GM/r \right)$$
$$+ v_{\parallel} B_{\perp}^2 / 8\pi - B_{\parallel} v_{\perp} B_{\perp} / 4\pi$$
at different locations
$$\Rightarrow \text{Heating}$$
Need *B* & *v*



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- 14:45-14:50 鈴木建 導入
- 14:50-15:05 松本琢磨 数値シミュレーションに基づく将来観測の展望
- 15:05-15:20 北川直優
  将来観測で期待される波動エネルギー流束の検出
- 15:20-15:35 末松芳法
  スピキュール観測で期待されるエネルギー流束の検出
- 15:35-15:45 飯田佑輔
  Hanle 効果による彩層中の波動検出の可能性
- 15:45-16:00 渡邊鉄哉
  Solar-C/EUVST における NeVII の観測
- 16:00-16:15 議論