# Ground-Based Perspectives of Solar-B Science and Operations



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- □ Solar-B is extremely exciting!
- When Solar-B becomes operational, my expectations will have been more than fulfilled!
- □ Thanks for involving us!



# **SOT = medium-sized ground-based telescope in space**

Interior structure and dynamics	Origin of solar activity cycle, dynamo	Origin of solar variability	Transient eruptions flares and coronal mass ejections	Heating of chromo- sphere and corona, origin of solar wind	Surface and atmo- sphere structure and dynamics	Exploring the unknown
S	noptic Gro	und-Based	<b>Felescopes</b>			
			Large Gro	und-Based	<b>Felescopes</b>	
		SOHO/S	DO			
				STEREO		
				TRACE		
				Solar-B		



## **Spatial Resolution/Field-of-View vs Wavelength Coverage**



![](_page_4_Picture_0.jpeg)

# SOT has a large field of view with high resolution

Solar-B will provide context images for the highest-resolution ground-based observations.

![](_page_4_Picture_3.jpeg)

#### March 30, 2001, W.C.Livingston

![](_page_5_Picture_0.jpeg)

# **On-Disk Coverage of Solar Atmosphere**

	temperature	velocity	magnetic field	
corona				
transition region				
chromosphere				
photosphere				

Ground-Based Optical

**Ground-Based Radio** 

Solar-B

![](_page_6_Picture_0.jpeg)

## Formation Heights (courtesy Han Uitenbroek)

![](_page_6_Figure_2.jpeg)

![](_page_7_Picture_1.jpeg)

- Think of Solar-B and ground-based telescopes as an observing system
- Select the combination of instruments in space and on the ground with the appropriate data analysis methods that have the highest likelihood of answering the scientific question
- Ground-based telescopes and Solar-B can have interchanging roles of supporting each other in terms of spatial coverage, temporal coverage, height coverage, and coverage of physical parameters
- Add *numerical simulations* to complement observing system

![](_page_8_Picture_0.jpeg)

- □ Solar-B launch around solar minimum, *think quiet sun*
- But, some 'theoretical active regions' (simple configurations)
- Magnetic field and temperature/density structures from the photosphere to the corona (canopies in quiet sun?)
- Magnetic field buffeting in photosphere (upward-traveling MHD waves?)
- Magnetic flux and field-strength evolution in the quiet sun (active-region leftovers, local dynamo?)

![](_page_9_Picture_1.jpeg)

- Need to schedule the observing system
- □ Agile scheduling that considers solar conditions
- Ground-based 'research' telescopes are much less flexible than SOHO, TRACE, Solar-B
- □ Change to ground-based scheduling
  - provide specific time for Solar-B collaboration
  - schedule it together with Solar-B scheduling and other groundbased facilities
  - accept proposals for these time slots together with Solar-B scheduling